Morphology and Topographic Profiles of Lobate Flow on Mars How convincing is the Water Ice Hypothesis?

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ABSTRACT

•Surface features observed East of the Hellas impact basin on Mars are indicative of viscous flow, and resemble Earth glaciers.

•This research compares the topographic profiles of the Martian flows to Terrestrial ice analogues.

•Terrestrial profiles are modeled using the equation of Glen's Flow Law.

•We seek to test whether ice is really the correct analogue as presumed by previous works.

METHODOLOGY

Martian topographic profiles are measured using the digital elevation model (DEM) obtained by the Mars Orbiter Laser Altimeter (MOLA).
Terrestrial equilibrium profiles are obtained by Glen's Flow Law.

 $\left(\frac{h}{H}\right)^{2+2/n} + \left(\frac{x}{L}\right)^{1+1/n} = 1$

•This equation expresses topographical elevation as a function of distance to some power n (where n = 3 for water-ice glaciers).

•This *n* exponent controls the plasticity of the material and is governed by stress and strain rate (Paterson 1994).

•An *n* value for mars different than 3 will suggest that ice is not a straightforward analogue.



GEOLOGIC SETTING

•Debris aprons are located mainly in the mountainous terrain on the eastern side of the Hellas basin.

•High-resolution images clearly indicate a material which has undergone viscous creep.



Image Credit: themis.asu.edu and http://cseligman.com

REFERENCES

•Peterson, W.S.B. The Physics of Glaciers, 3rd ed. 1994. •Li, H. et al. Maritan Southern Hemisphere Debris Aprons. Lunar and Planetary Science. 2006.

•Head, J.W. et al. Tropical to mid-latitude snow and ice accumulation, flow, and glaciation on mars. Nature. 2005.

RESULTS

•Below: Fourteen Martian topological profiles overlapped by the Glen's Flow Law, with *n* value plotted at n = 1 (green), n = 3 (blue), and n = 20 (cyan). Image also shown of approximate transect lines. •Profiles clearly do not correlate with any



CONCLUSIONS AND IMPLICATIONS

•These finding indicate the material flowing is either not ice or is not dictated by the properties of Glen's Flow Law.

•Re- examination of the many assumptions within the Glen's Flow Law equation. What are some contributing feedback processes such as sublimination, bedrock topography, temperature effects, etc.

•Other potential analogues? Previous published work suggests the Martian flows are water-ice glaciers. Recent discovery of abundant salts on Mars invite this hypothesis to be reconsidered. Another possible terrestrial analogue may be the salt glaciers of Iran.